Space Studies of the Upper Atmospheres of the Earth and Planets including Reference Atmospheres (C)

Planetary Upper Atmospheres, Ionospheres and Magnetospheres (C32)

THE STRUCTURE OF THE MARS IONOSPHERE

Dr. Martin Paetzold, mpaetzol@uni-koeln.de Rhenish Institute for Environmental Research at the University of Cologne, Cologne, Germany Silvia Tellmann, stellman@uni-koeln.de Rhenish Institute for Environmental Research at the University of Cologne, Cologne, Germany Kerstin Peter, kerstin.peter@uni-koeln.de Rhenish Institute for Environmental Research at the University of Cologne, Cologne, Germany Michael Mendillo, mendillo@bu.edu Boston University, Boston, Massachusetts, United States Paul Withers, withers@bu.edu Boston University, Boston, Massachusetts, United States Bernd Haeusler, Bernd.Haeusler@unibw-muenchen.de Bundeswehr University, Munich, Munich, Germany David P. Hinson, dhinson@stanford.edu Stanford University, Stanford, California, United States G.L. Tyler, len.tyler@stanford.edu Stanford University, Stanford, California, United States

The Mars Express Radio Science Experiment MaRS sounds the ionosphere of Mars at microwavelengths and covers altitudes from the base of the ionosphere at 80 km to the ionopause at altitudes between 300 km and 800 km. The Mars ionosphere consists of a lower secondary layer M1 at about 110 km, and the main layer M2 at about 135 km altitude, both formed by solar radiation at X-ray and EUV, respectively. The precise and detailed observations of the Mars Express radio science experiment indicates the presence of another layer M3 in the topside above the main layer M2 with a shape of a Chapman function as the transition region between the photochemically induced and Chapman-like M1 and M2 layers and the transport dominated highly dynamical topside region above 200 km altitude. Sporadically, a region of enhanced ionisation below the M1 layer can be observed which is caused by the infall of meteors into the atmosphere.