

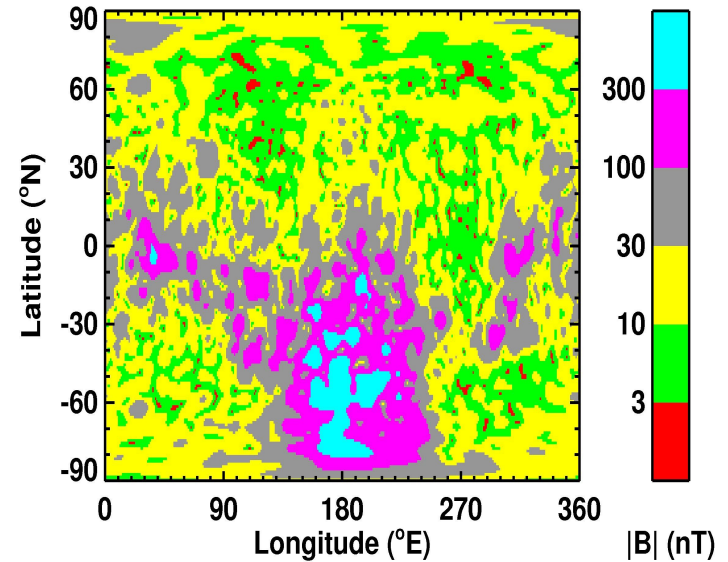
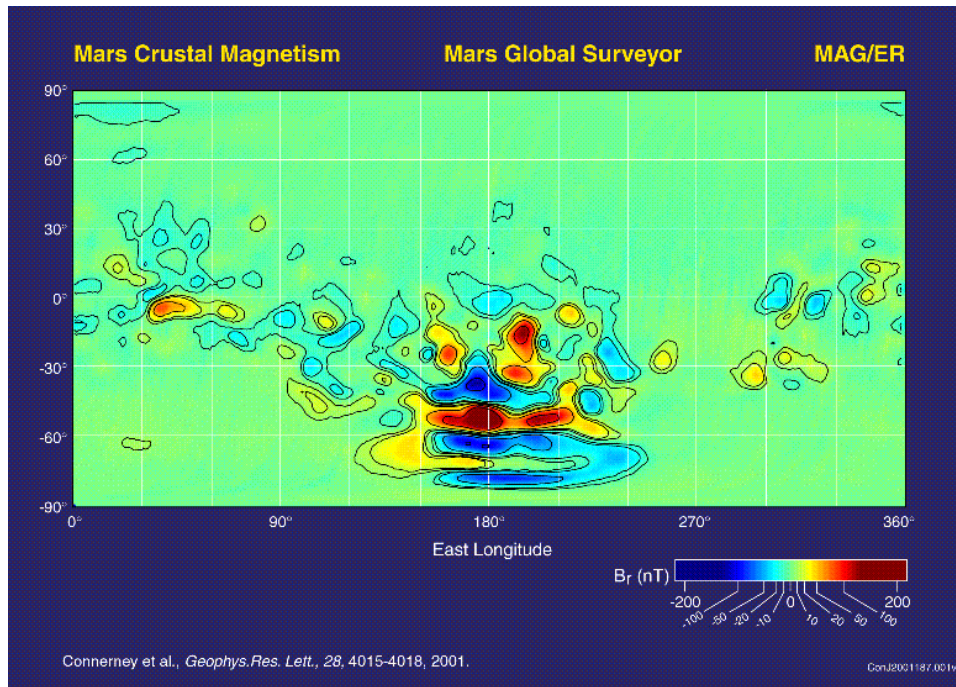


The unusual electrodynamics of the ionosphere of Mars

Paul Withers, **Majd Matta**, Michael
Mendillo
Boston University

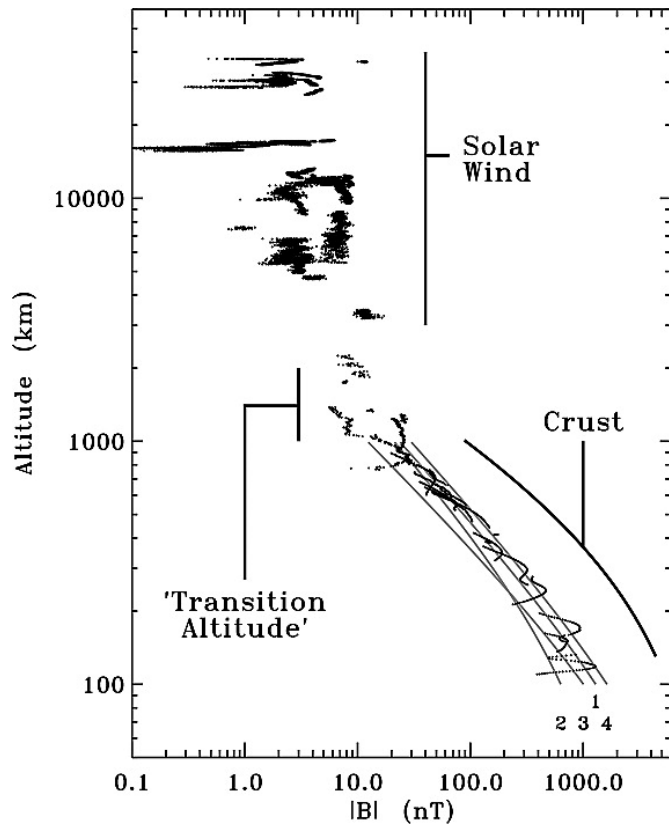
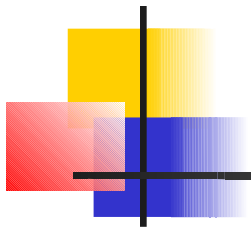
ESPC 2010, Sept 19-24, Rome, Italy

Martian Patchy Fields

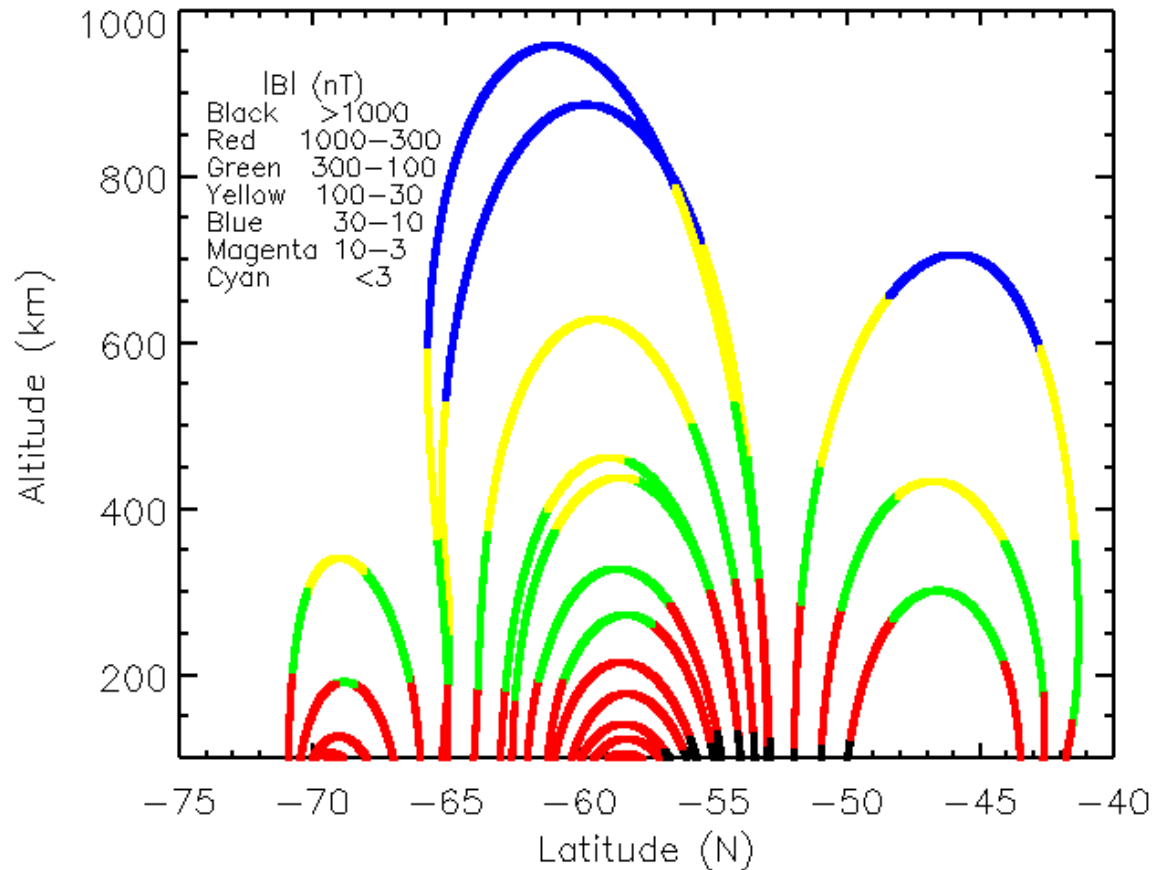


Arkani-Hamed model

Martian Field Lines

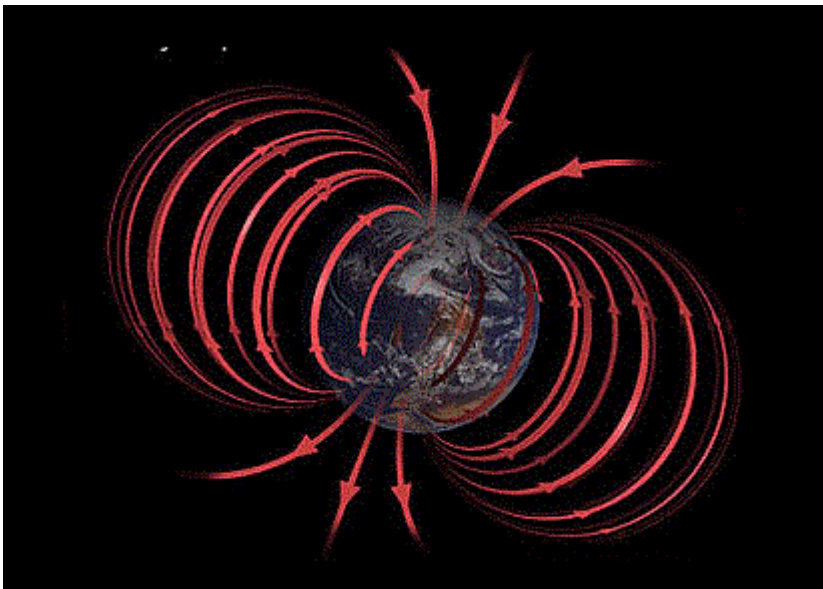


Brain et al. 2003

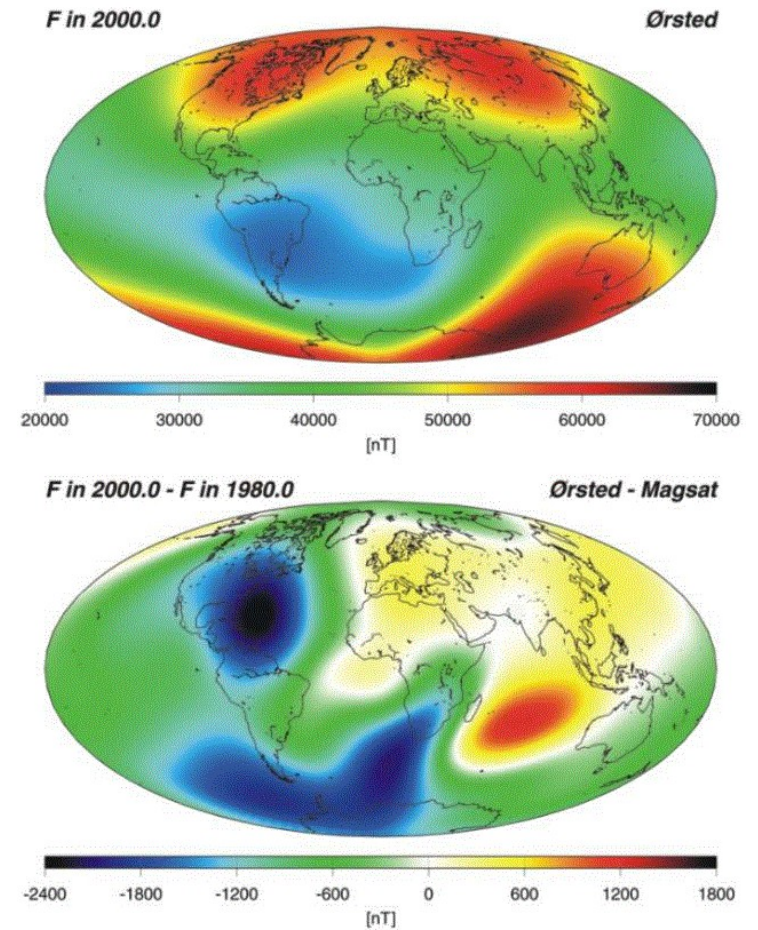


Mendillo and Withers, 2008

Earth's Field Lines

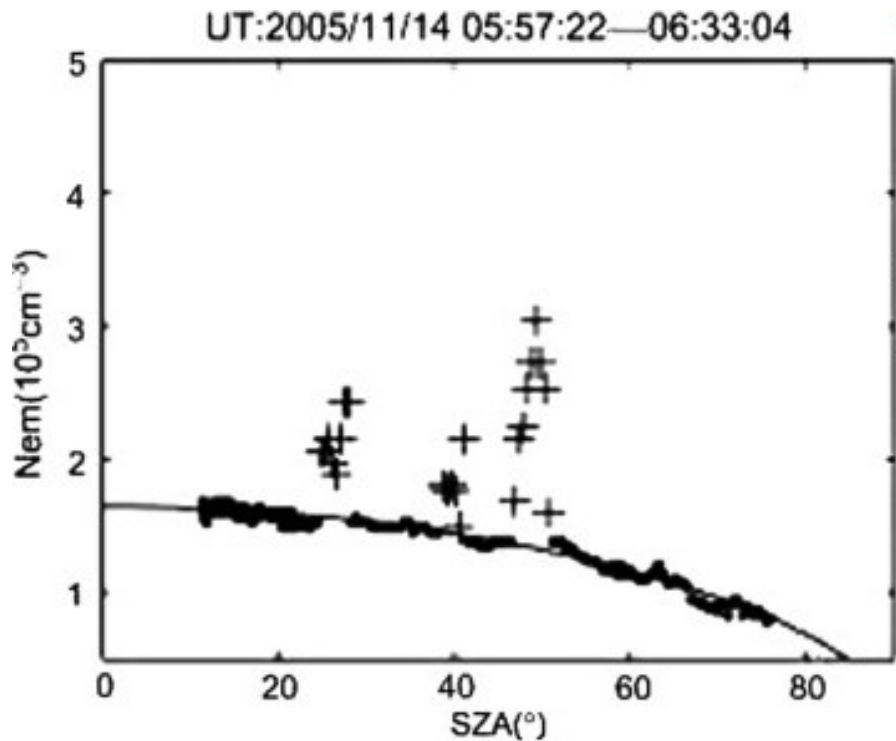
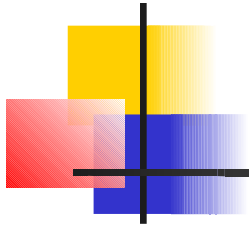


http://www.windows2universe.org/glossary/particle_motion.html

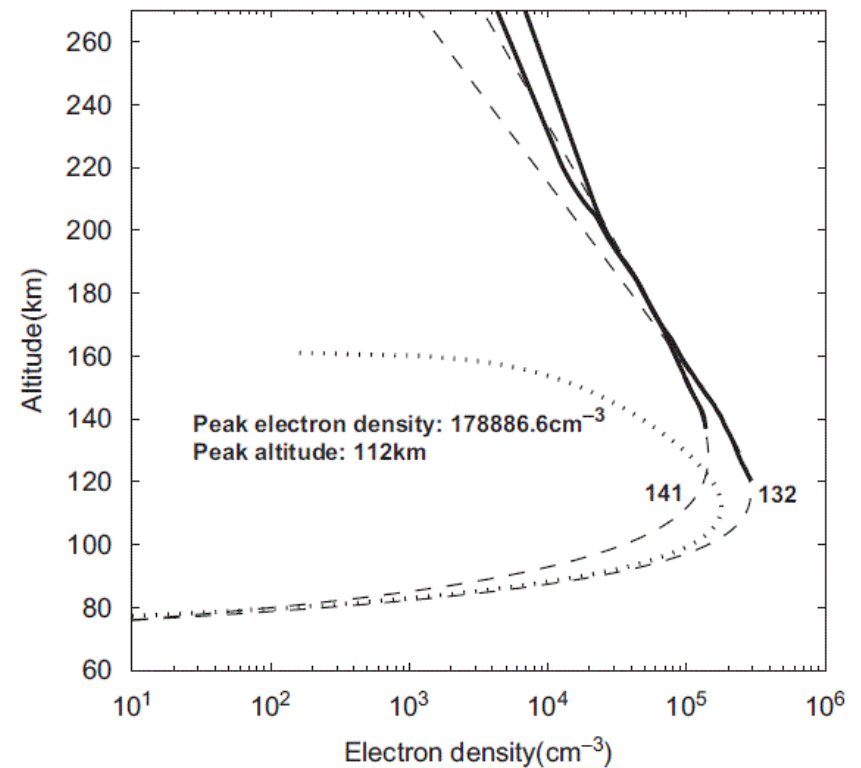


http://core2.gsfc.nasa.gov/terr_mag/core.html

Effect of Crustal B on Ionosphere

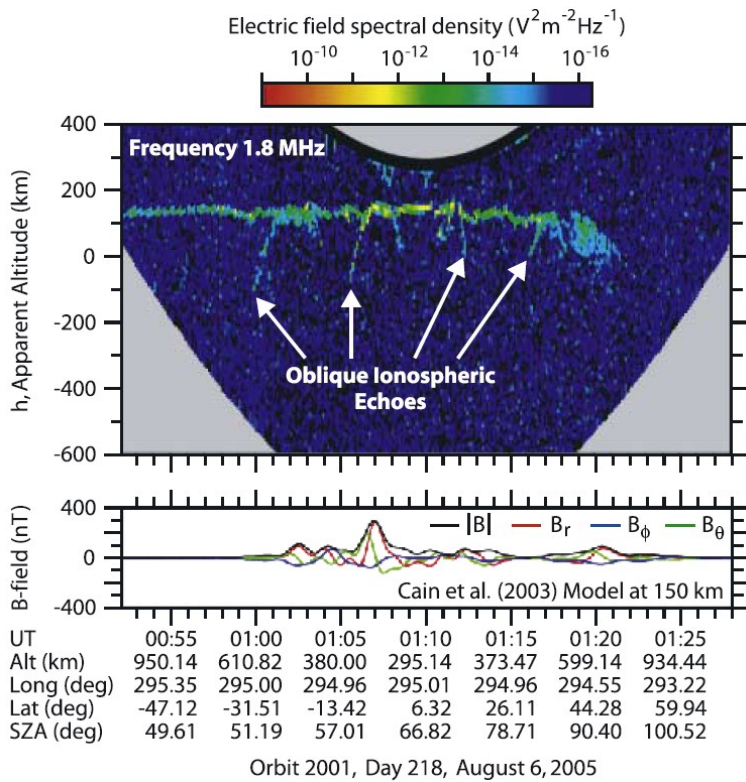


Nielson et al. 2007a, from Fig 1

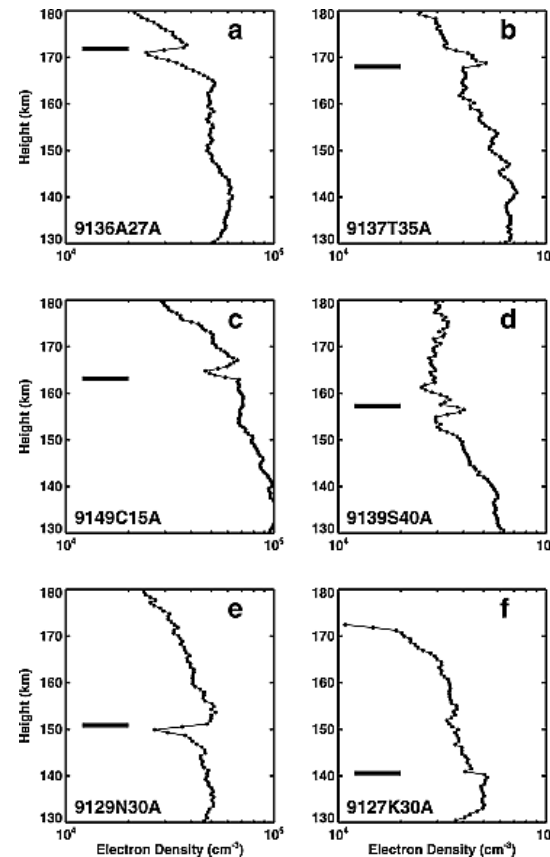


Nielson et al. 2007a Fig 5

More effects of Crustal B



Duru et al., 2006, Fig 4



- Ion densities
- Electron Densities
- Temperatures
- Neutral Atmosphere

Withers 2005a, Fig 2

Theory (Withers, 2008)

$$\frac{\partial N_j}{\partial t} + \underline{\nabla} \cdot (N_j \underline{v}_j) = P_j - L_j$$

$$0 = m_j \underline{g} - \frac{1}{N_j} \underline{\nabla} (N_j k T_j) + q_j \underline{E}' + q_j B \underline{\underline{\Lambda}} \underline{w}_j - m_j \nu_{jn} \underline{w}_j$$

$$\underline{w}_j = \left(m_j \nu_{jn} \underline{\underline{I}} - q_j B \underline{\underline{\Lambda}} \right)^{-1} \left(m_j \underline{g} - \frac{1}{N_j} \underline{\nabla} (N_j k T_j) + q_j \underline{E}' \right)$$

$$\underline{w}_j = \frac{1}{m_j \nu_{jn}} \left(\underline{\underline{I}} - \kappa_j \underline{\underline{\Lambda}} \right)^{-1} \left(m_j \underline{g} - \frac{1}{N_j} \underline{\nabla} (N_j k T_j) + q_j \underline{E}' \right)$$

$$\underline{w}_j = \frac{1}{N_j q_j} \left(\underline{\underline{Q}}_j + \underline{\underline{S}}_j \underline{E}' \right)$$

$$\underline{J} = \sum_j N_j q_j \underline{w}_j \quad \longrightarrow \quad \underline{J} = \underline{\underline{Q}} + \underline{\underline{S}} \underline{E}'$$

Traditional Conductivity Tensor

$$\underline{J} = \underline{\underline{\sigma}} \underline{E}' , \text{ where } \underline{\underline{\sigma}} = \begin{pmatrix} \sigma_P & -\sigma_H & 0 \\ \sigma_H & \sigma_P & 0 \\ 0 & 0 & \sigma_0 \end{pmatrix}$$

$$\sigma_P = \sum_{j=1}^M \frac{N_j q_j}{B} \frac{v_j / \omega_j}{1 + (v_j / \omega_j)^2}$$

$$\sigma_H = \sum_{j=1}^M \frac{N_j q_j}{B} \frac{1}{1 + (v_j / \omega_j)^2}$$

$$\sigma_0 = \sum_{j=1}^M \frac{N_j q_j}{B} \frac{\omega_j}{v_j}$$

(Forbes, 1981)

General Conductivity

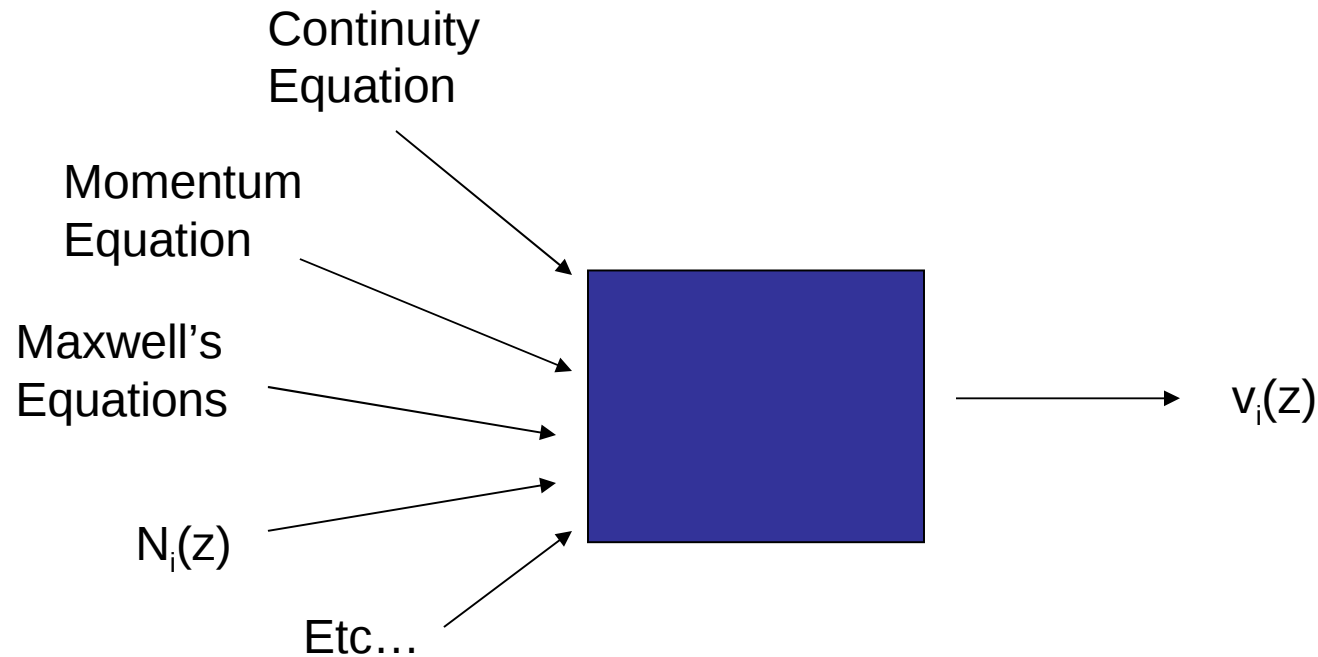
Tensor:

$$\underline{\underline{S}} = \sum_j \frac{N_j q_j^2}{m_j \nu_{jn}} \begin{pmatrix} \frac{1}{(1 + \kappa_j^2)} & \frac{\kappa_j}{(1 + \kappa_j^2)} & 0 \\ -\kappa_j & 1 & 0 \\ \frac{1}{(1 + \kappa_j^2)} & \frac{1}{(1 + \kappa_j^2)} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

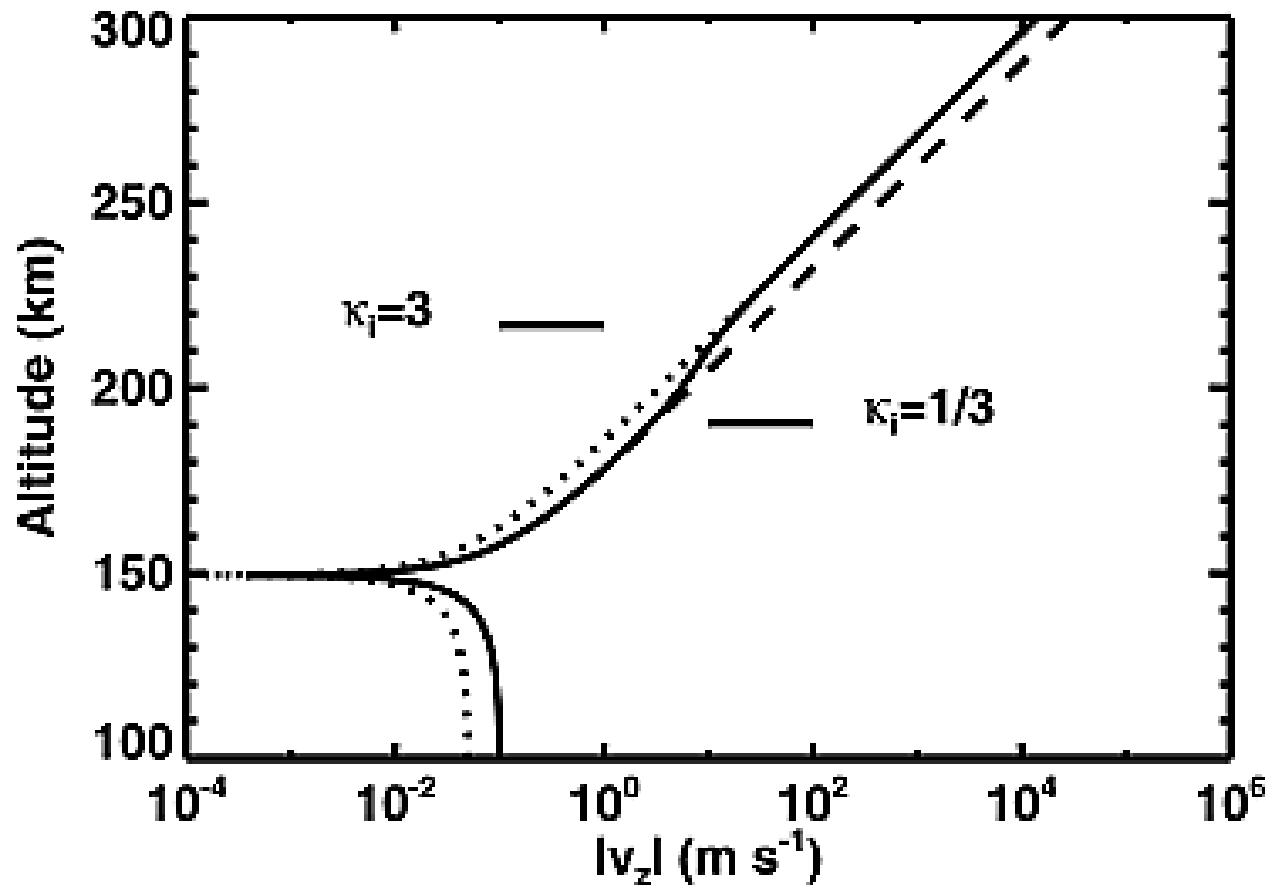
(Withers, 2008)



Determining Ion Velocities

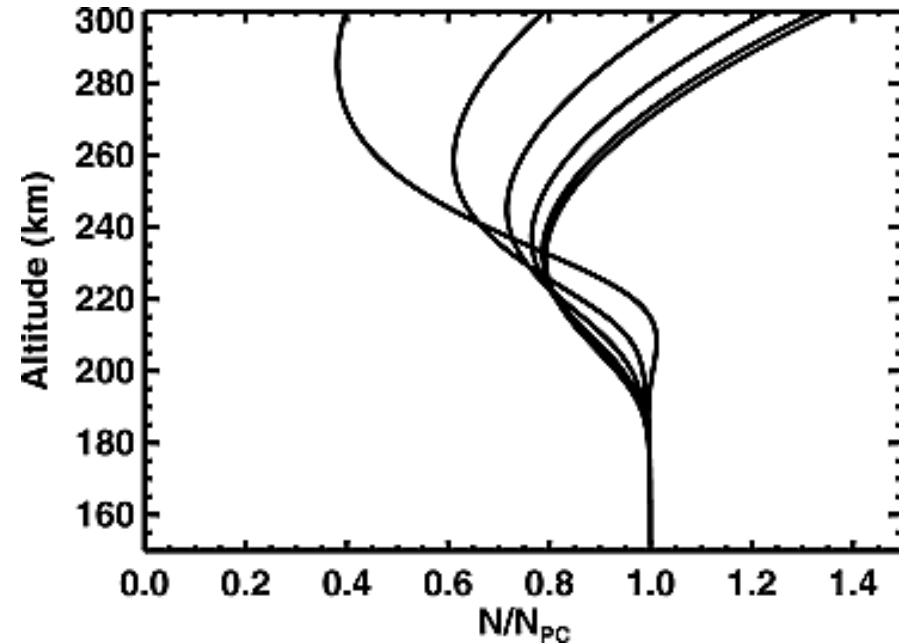
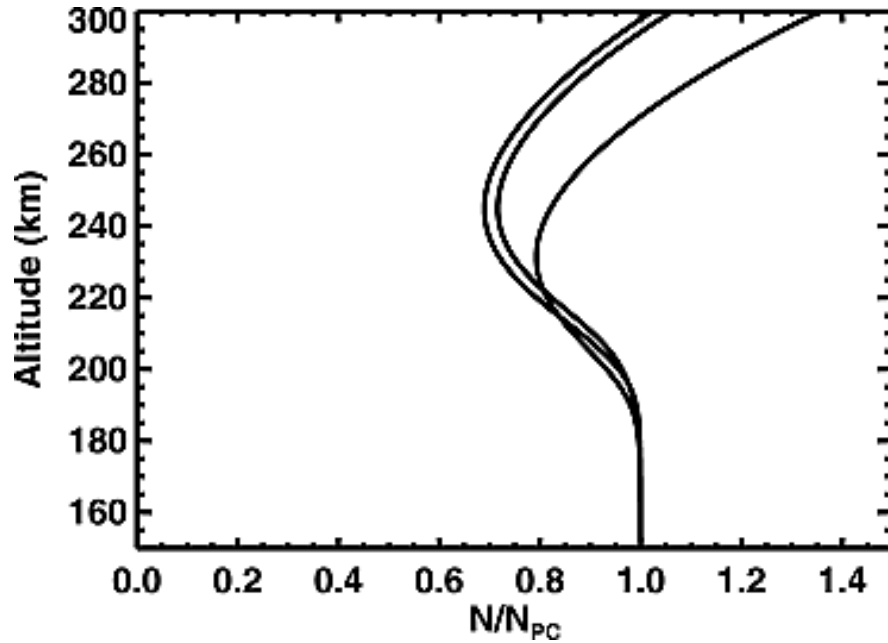
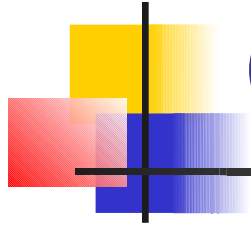


More Theory



(Withers, 2008)

Application of General Conductivity



(Withers, 2008: Figures 7 and 9)



Next, a 2D Model:

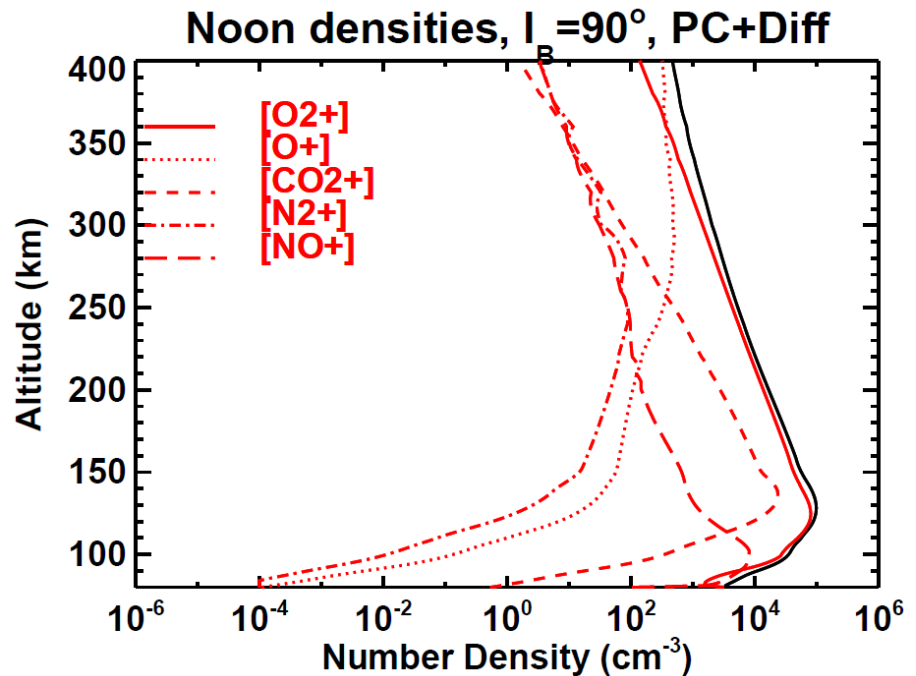
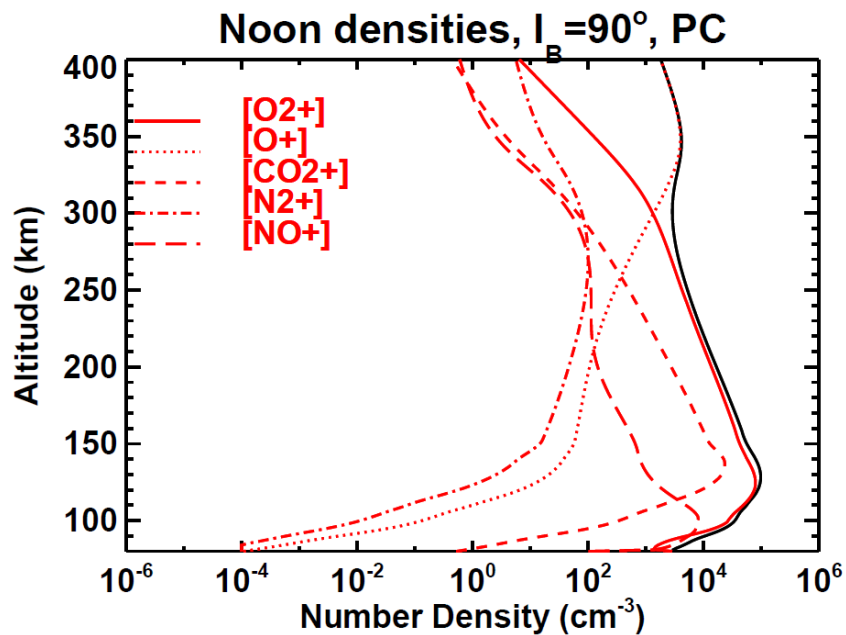
- Objective is to model, as accurately as possible, the generation of 5 major ions in the Martian ionosphere between 80-400 km.
- Then to analyze this model as input parameters change (location over surface, surrounding magnetic environment, etc...)



2D Model Methodology

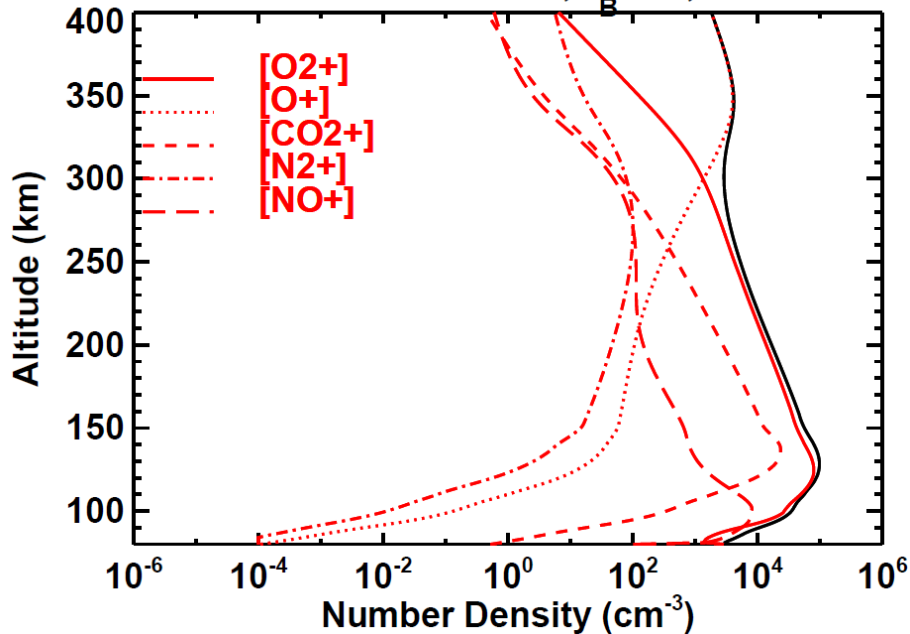
- PC Production
- PC Loss
- Transport
- Calculation of 2D ion velocities

2D Model, Preliminary Results

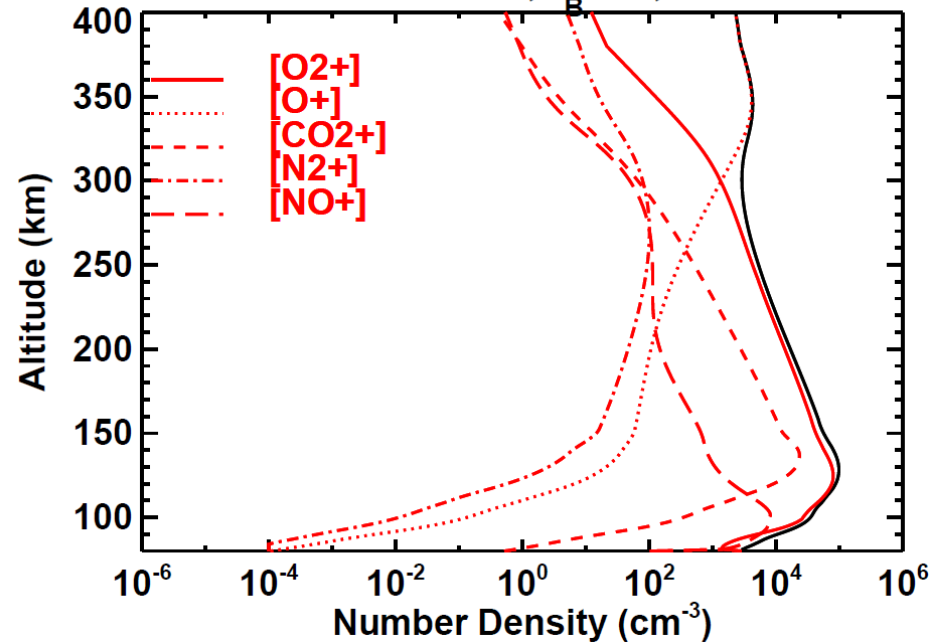


More Preliminaries

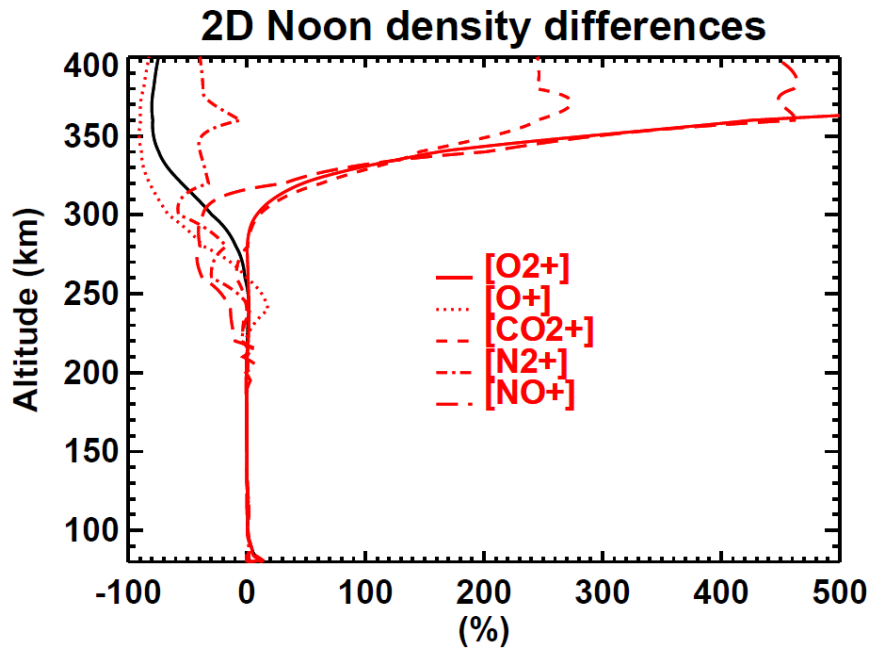
Noon densities, $I_B = 0^\circ$, PC



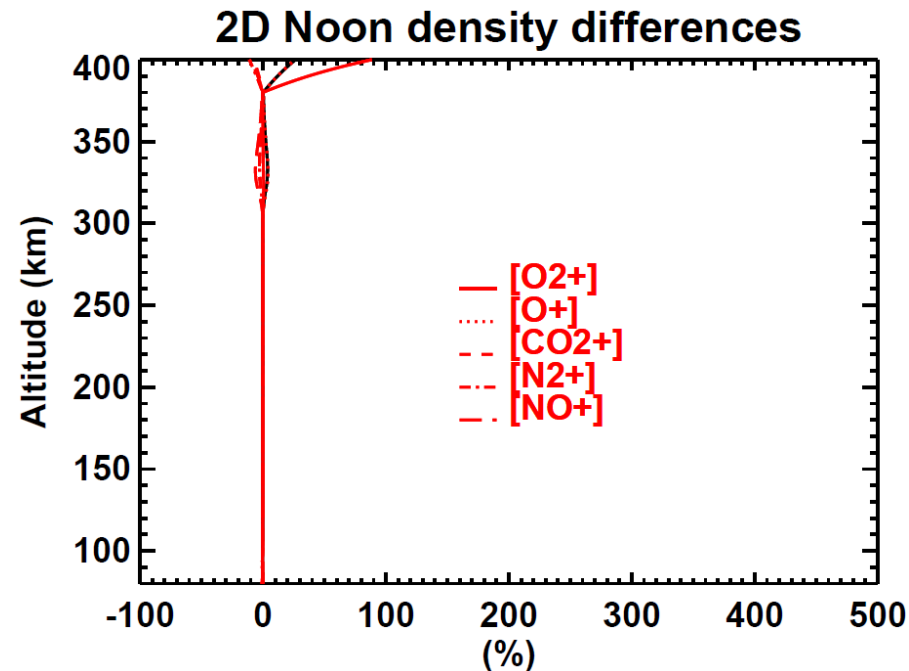
Noon densities, $I_B = 0^\circ$, PC+Diff



More 2D Model Preliminaries



Difference between PC only and PC with Diffusion densities for $I_B=90^\circ$



Difference between PC only and PC with Diffusion densities for $I_B=0^\circ$



Conclusions

- Martian crustal fields generate effects on the ionosphere that have been measured and modeled (1D).
- The most general representations of underlying physics apply to the regions of Martian ionospheres that should not be generalized for terrestrial-like cases.
- Further study and modeling (2D) is required to improve our understanding of the phenomena governing such regions.